

On the integrable deformations of continuous-time dynamical systems

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A **deformation** of the system

$$\dot{\mathbf{x}} = \mathbf{f}(\mathbf{x}), \quad \mathbf{x} = (x, y, z), \quad \mathbf{f} = (f_1, f_2, f_3) \quad (1)$$

is given by a system of the form

$$\begin{cases} \dot{x} = f_1(x, y, z) + u_1(x, y, z, \mathbf{d}) \\ \dot{y} = f_2(x, y, z) + u_2(x, y, z, \mathbf{d}) \\ \dot{z} = f_3(x, y, z) + u_3(x, y, z, \mathbf{d}) \end{cases}, \quad (2)$$

where the vector \mathbf{d} contains the deformation parameters such that

$$\text{if } \mathbf{d} \rightarrow \mathbf{0} \text{ then } (2) \rightarrow (1).$$

The functions added to a system by deformation can be viewed as control functions and as a consequence some natural queries arise:

- how does the stability change
 - can be stabilized some states/orbits
 - how is the periodic motion affected
 - can be obtained new chaotic systems
 - can be stabilized some chaotic trajectories
- etc.

A function $F = F(x, y, z)$ is a **constant of motion** of the system

$$\dot{\mathbf{x}} = \mathbf{f}(\mathbf{x}), \quad \mathbf{x} = (x, y, z), \quad \mathbf{f} = (f_1, f_2, f_3)$$

if

$$\dot{F} = \nabla F \cdot \mathbf{f} = 0.$$

The above system is **integrable** if there are two independent functions that are constants of motion of this system. In fact, such a system is a Hamilton-Poisson system.

A deformation of this system is an **integrable deformation** if it has two independent constants of motion too.

Integrable deformations for a three-dimensional Hamilton-Poisson system

Let H, C be independent constants of motion of the system

$$\dot{\mathbf{x}} = \mathbf{g}(\mathbf{x}), \quad \mathbf{x} = (x, y, z), \quad \mathbf{g} = (g_1, g_2, g_3). \quad (3)$$

It is known there is a function $\mu \in C^1$ (see, for example, [1]) such that

$$\mathbf{g} = \mu \nabla H \times \nabla C.$$

Indeed, because H and C are constants of motion it follows

$$\nabla H \cdot \mathbf{g} = 0, \nabla C \cdot \mathbf{g} = 0 \Leftrightarrow \nabla H \perp \mathbf{g}, \nabla C \perp \mathbf{g}.$$

Therefore $\nabla H \times \nabla C$ and \mathbf{g} are parallel, whence the conclusion follows.

¹M. Gürses, G. S. Guseinov, K. Zheltukhin, *Dynamical systems and Poisson structures*, J. Math. Phys. 50 (2009) 112703.

Now, we consider the system

$$\dot{\mathbf{x}} = \mu \nabla \tilde{H} \times \nabla \tilde{C}, \quad \tilde{H}(\mathbf{x}) = H(\mathbf{x}) + k_1 u(\mathbf{x}), \quad \tilde{C}(\mathbf{x}) = C(\mathbf{x}) + k_2 v(\mathbf{x}), \quad (4)$$

where u, v are arbitrary differentiable functions and $k_1, k_2 \in \mathbb{R}$.

An easy computation shows that system (4) becomes

$$\dot{\mathbf{x}} = \mu \nabla H \times \nabla C + k_2 \mu \nabla H \times \nabla v + k_1 \mu \nabla u \times \nabla C + k_1 k_2 \mu \nabla u \times \nabla v. \quad (5)$$

Remark. Because \tilde{H} and \tilde{C} are constants of motion of system (5), it is clear that this system is an integrable deformation of the integrable system $\dot{\mathbf{x}} = \mu \nabla H \times \nabla C$.

Consider the motion of a dynamical system

$$\dot{\mathbf{x}} = \mathbf{f}(\mathbf{x}), \quad \mathbf{x} = (x, y, z), \quad \mathbf{f} = (f_1, f_2, f_3)$$

in the phase space as the flow of a fluid. Hence the logarithmic rate of the volume change is expressed as²

$$\frac{1}{V} \frac{dV}{dt} = \nabla \cdot \mathbf{f},$$

where

$$\nabla \cdot \mathbf{f} = \frac{\partial f_1}{\partial x} + \frac{\partial f_2}{\partial y} + \frac{\partial f_3}{\partial z}$$

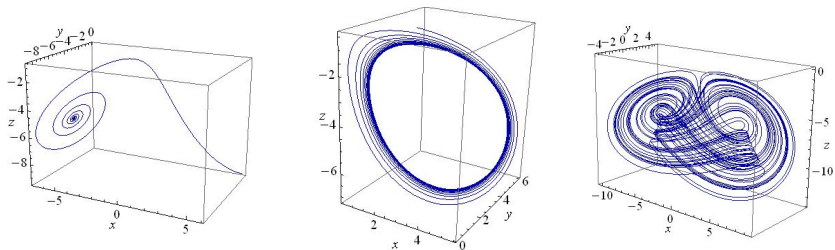
is the divergence of the vector field \mathbf{f} .

If the divergence is less than zero, then the system is called **dissipative**.

²E. N. Lorenz, Deterministic nonperiodic flow, J. Atmos. Sci. 20 (1963) 130–141.

A dissipative system has "a state space volume that decreases on average along the trajectory so that the orbit approaches an attractor of measure zero in the state space"³.

For a three-dimensional dissipative dynamical system these attractors can be stable foci, limit cycles or strange chaotic attractors, and they can coexist.⁴



³J. C. Sprott, A dynamical system with a strange attractor and invariant tori, Phys. Lett. A 378 (2014) 1361–1363.

⁴J. C. Sprott, X. Wang, G. Chen, Coexistence of point, periodic and strange attractors, Int. J. Bifurcation and Chaos 23 (2013) 1350093–1–5.

The integrable deformations method for a three-dimensional system of differential equations

Let $\Omega \subseteq \mathbb{R}^3$ be an open set. We consider the following three-dimensional dynamical system on Ω

$$\dot{\mathbf{x}} = \mathbf{f}(\mathbf{x}) , \quad \mathbf{x} = (x, y, z), \mathbf{f} = (f_1, f_2, f_3). \quad (6)$$

We assume that system (6) has a Hamilton-Poisson part described by a vector field \mathbf{g} (see, for example, Kolmogorov systems⁵). Therefore system (6) can be written in the form

$$\dot{\mathbf{x}} = \mathbf{g}(\mathbf{x}) + \mathbf{h}(\mathbf{x}) , \quad (7)$$

where the system

$$\dot{\mathbf{x}} = \mathbf{g}(\mathbf{x}) \quad (8)$$

has two functionally independent constants of motion, $H = H(\mathbf{x})$ and $C = C(\mathbf{x})$.

⁵V. I. Arnold, *Kolmogorov's hydrodynamic attractors*, Proc. Roy. Soc. Lond. A 434 (1991) 19–22.

As above, we construct integrable deformations of the Hamilton-Poisson part $\dot{\mathbf{x}} = \mu \nabla H \times \nabla C$ taking $\tilde{H} = H + \alpha$, $\tilde{C} = C + \beta$. Then, adding the non-conservative part \mathbf{h} , we obtain the integrable deformations of the given system⁶ :

$$\dot{\mathbf{x}} = \mathbf{f}(\mathbf{x}) + \mu \nabla H \times \nabla \beta + \mu \nabla \alpha \times \nabla C + \mu \nabla \alpha \times \nabla \beta. \quad (9)$$

Remark. If the function μ is constant, then the divergence of system (9) is equal to the divergence of the initial system (6), $\dot{\mathbf{x}} = \mathbf{f}(\mathbf{x})$.

⁶C. Lăzureanu, *Integrable Deformations of Three-Dimensional Chaotic Systems*, Int. J. Bifurcat. Chaos 28 (5) (2018) 1850066.

- there is an integrable deformation of a given chaotic system that has also a chaotic behavior
- there is an integrable deformation of a nonchaotic version of a chaotic system that presents a chaotic attractor
- there is an integrable deformation which stabilizes an unstable state of a given system

Integrable deformations of the Lorenz system

The first chaotic attractor in a three-dimensional autonomous system is given by the Lorenz system⁷

$$\begin{cases} \dot{x} &= a(y - x) \\ \dot{y} &= cx - y - xz \\ \dot{z} &= -bz + xy \end{cases}, \quad (a = 10, b = \frac{8}{3}, c = 28). \quad (10)$$

We consider the following Hamilton-Poisson part of system (10)

$$\begin{cases} \dot{x} &= 0 \\ \dot{y} &= cx - xz \\ \dot{z} &= xy \end{cases}, \quad (11)$$

with the constants of motion $H = \frac{1}{2}x^2$, $C = \frac{1}{2}y^2 + \frac{1}{2}z^2 - cz$.

System (11) writes $\dot{\mathbf{x}} = \nabla H \times \nabla C$.

⁷E. N. Lorenz, *Deterministic nonperiodic flow*, J. Atmos. Sci. 20 (2) (1963) 130–141.

Therefore we give the following integrable deformation of Lorenz system⁸

$$\begin{cases} \dot{x} &= a(y - x) + (z - c)\alpha_y - y\alpha_z + \alpha_y\beta_z - \alpha_z\beta_y \\ \dot{y} &= cx - y - xz - x\beta_z - (z - c)\alpha_x - \alpha_x\beta_z + \alpha_z\beta_x \\ \dot{z} &= -bz + xy + x\beta_y + y\alpha_x + \alpha_x\beta_y - \alpha_y\beta_x \end{cases} \quad (12)$$

1. We deform a chaotic system and obtain another chaotic system.

Choosing the deformation functions $\alpha = \frac{g}{2}z^2$, $\beta = 0$, we obtain the system

$$\begin{cases} \dot{x} &= a(y - x) - gyz \\ \dot{y} &= cx - y - xz \\ \dot{z} &= -bz + xy \end{cases}, \quad (13)$$

where $g \in \mathbb{R}$ is the deformation parameter. We notice that if $g = 1$, then this system is so-called Qi system⁹, which is a chaotic system.

⁸C. Lăzureanu, *Integrable Deformations of Three-Dimensional Chaotic Systems*, Int. J. Bifurcat. Chaos Int. J. Bifurcat. Chaos 28 (5) (2018) 1850066.

⁹G. Qi, G. Chen, S. Du, Z. Chen, Z. Yuan, *Analysis of a new chaotic system*, Physica A 352 (2005) 295–308.

2. We deform a non-chaotic version of a chaotic system and obtain a chaotic system.

Setting $\alpha = 0$, $\beta = gxy$, system (12) becomes

$$\begin{cases} \dot{x} &= a(y - x) \\ \dot{y} &= cx - y - xz \\ \dot{z} &= -bz + xy + gx^2 \end{cases} . \quad (14)$$

The Lorenz system is not a chaotic one for some values of the parameters a, b, c . However, for the same values of a, b, c , we can choose deformation functions such that the obtained system has a chaotic behavior. For example, setting $a = 10, b = 5, c = 28$, the Lorenz attractor (10) is an asymptotically stable equilibrium point, and, if $g = 3$, the attractor of the integrable deformation of Lorenz system given by (14) is chaotic.

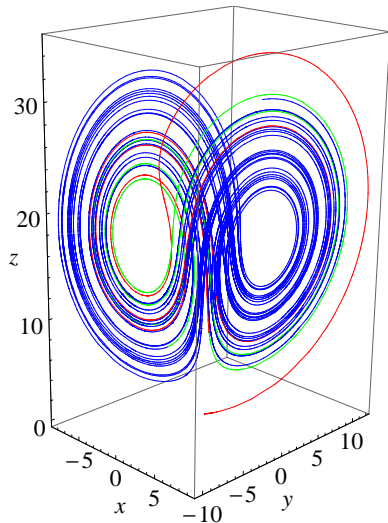
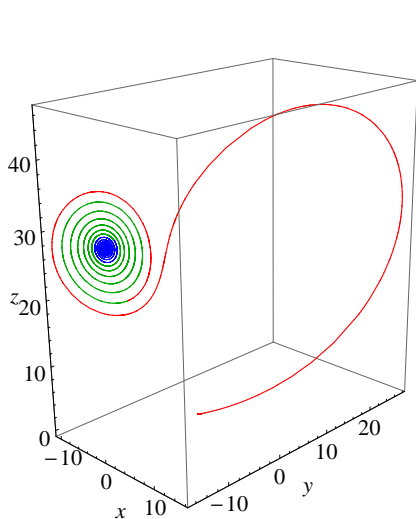


Figure: The Lorenz attractor (left) and one of its deformations (right, $g = 3$), for $a = 10, b = 5, c = 28$ ($t \in [0, 2]$ blue line, $t \in [2, 6]$ green line; initial conditions $x(0) = y(0) = z(0) = 0.1$)

Integrable deformations of Rössler system

In 1976, Rössler¹⁰ proposed the following “prototype equation to the Lorenz model of turbulence which contains just one (second-order) nonlinearity”

$$\begin{cases} \dot{x} &= -y - z \\ \dot{y} &= x + ay \\ \dot{z} &= -cz + xz + b \end{cases}, \text{ where } a = 0.2, b = 0.2, c = 5.7. \quad (15)$$

We consider the following Hamilton-Poisson part of system (15)

$$\begin{cases} \dot{x} &= -y \\ \dot{y} &= x \\ \dot{z} &= 0 \end{cases}, \quad (16)$$

with the constants of motion $H = z$, $C = \frac{1}{2}x^2 + \frac{1}{2}y^2$. We get that $\mu = 1$.

¹⁰O. E. Rössler, *An equation for continuous chaos*, Phys. Lett. A 57 (1976) 397–398.

We give the following integrable deformation of Rössler system¹¹

$$\begin{cases} \dot{x} &= -y - z - \beta_y - y\alpha_z + \alpha_y\beta_z - \alpha_z\beta_y \\ \dot{y} &= x + ay + \beta_x + x\alpha_z - \alpha_x\beta_z + \alpha_z\beta_x \\ \dot{z} &= -cz + xz + b + y\alpha_x - x\alpha_y + \alpha_x\beta_y - \alpha_y\beta_x \end{cases} . \quad (17)$$

Considering the deformation functions $\alpha = \frac{g}{2}z^2$, $\beta = 0$, system (17) becomes

$$\begin{cases} \dot{x} &= -y - z - gyz \\ \dot{y} &= x + ay + gxz \\ \dot{z} &= -cz + xz + b \end{cases} , \quad (18)$$

where the parameters a, b, c take the above-mentioned values, and $g \in \mathbb{R}$ is the deformation parameter.

¹¹C. Lăzureanu, *Integrable Deformations of Three-Dimensional Chaotic Systems*, Int. J. Bifurcat. Chaos Int. J. Bifurcat. Chaos 28 (5) (2018) 1850066.

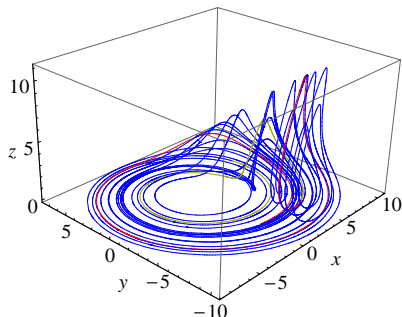
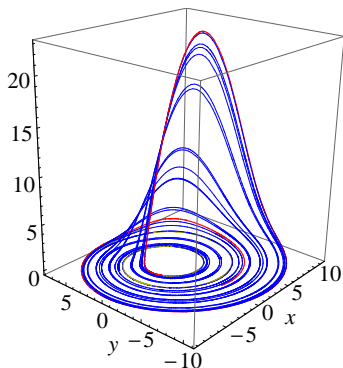


Figure: The Rössler attractor (left) and one of its deformations (right, $g = -0.25$), $a = 0.2$, $b = 0.2$, $c = 5.7$ ($t \in [0, 10]$ blue line, $t \in [10, 20]$ green line; initial conditions $x(0) = 0, y(0) = -6.78, z(0) = 0.02$)

A generalization: Integrable Deformations of the Maximally Superintegrable Systems

In classical mechanics, a superintegrable system on a $2n$ -dimensional phase space is a completely integrable Hamiltonian system which possesses more functionally independent first integrals than degrees of freedom. Moreover, such a system is called maximally superintegrable¹² if the number of the independent first integrals is $2n - 1$.

Similarly, for an arbitrary natural number n , a system of first-order differential equations on \mathbb{R}^n which has $n - 1$ functionally independent constants of motion is called a maximally superintegrable system.

If we can identify a maximally superintegrable part of the considered system, then we proceed as in the three-dimensional case.

¹²A.V. Tsiganov, *On maximally superintegrable systems*, Regul. Chaotic Dyn. 13 (2008) 178–190.

The functions C_1, C_2, \dots, C_{n-1} define the maximally superintegrable Hamilton-Poisson system

$$\dot{x}_k = \{H, x_k\}_{C_1, \dots, C_{n-2}}^\nu, \quad k \in \{1, 2, \dots, n\}, \quad (19)$$

with the Hamiltonian $H := C_{n-1}$ and the Poisson bracket¹³ given by

$$\{f, g\}_{C_1, \dots, C_{n-2}}^\nu := \nu \cdot \frac{\partial(C_1, C_2, \dots, C_{n-2}, f, g)}{\partial(x_1, x_2, \dots, x_{n-1}, x_n)}, \quad f, g \in C^\infty(\mathbb{R}^n, \mathbb{R}). \quad (20)$$

Conversely, for a given maximally superintegrable system, there is a unique function ν such that the system takes the form (19).

¹³R.M. Tudoran, *A normal form of completely integrable systems*, J. Geom. Phys. 62 (2012) 1167–1174.

Let $\alpha_1, \alpha_2, \dots, \alpha_{n-1} \in C^1(\Omega, \mathbb{R})$ be arbitrary functions such that the functions

$$I_k = C_k + g_k \alpha_k, \quad k \in \{1, 2, \dots, n-1\}$$

are functionally independent on an open set $\Omega \subseteq \mathbb{R}^n$, where $g_1, g_2, \dots, g_{n-1} \in \mathbb{R}$. Then an integrable deformation of the above system is given by¹⁴

$$\begin{aligned} \dot{x}_k &= \{C_{n-1}, x_k\}_{C_1, \dots, C_{n-2}}^\nu + g_1 \{C_{n-1}, x_k\}_{\alpha_1, C_2, \dots, C_{n-2}}^\nu \\ &+ \sum_{i=2}^{n-3} g_i \{C_{n-1}, x_k\}_{I_1, \dots, I_{i-1}, \alpha_i, C_{i+1}, \dots, C_{n-2}}^\nu \\ &+ g_{n-2} \{C_{n-1}, x_k\}_{I_1, \dots, I_{n-3}, \alpha_{n-2}}^\nu + g_{n-1} \{\alpha_{n-1}, x_k\}_{I_1, \dots, I_{n-2}}^\nu, \end{aligned}$$

$$k \in \{1, 2, \dots, n\}.$$

¹⁴C. Lăzureanu, *On the Integrable Deformations of the Maximally Superintegrable Systems*, Symmetry 13 (2021) 1000.

Consider a Lotka-Volterra type system¹⁵ given by

$$\dot{x}_1 = x_1 x_2, \quad \dot{x}_2 = x_2(x_3 - x_1), \quad \dot{x}_3 = x_3(x_4 - x_2), \quad \dot{x}_4 = -x_3 x_4,$$

with a constant of motion given by $H = x_1 + x_2 + x_3 + x_4$.

Apparently, this system is not maximally superintegrable. However, we can consider a maximally superintegrable part of it, namely

$$\dot{x}_1 = 0, \quad \dot{x}_2 = x_2 x_3, \quad \dot{x}_3 = -x_2 x_3, \quad \dot{x}_4 = -x_3 x_4.$$

Indeed, the functions $C_1 = x_1$, $C_2 = x_2 + x_3$, and $C_3 = x_2 x_4$ are constants of motion. Moreover, the above system takes the form

$$\dot{x}_k = \{C_3, x_k\}_{C_1, C_2}^\nu, \quad k \in \{1, 2, 3, 4\},$$

where $\nu = x_3$.

Now we can construct integrable deformations.

¹⁵O.I. Bogoyavlenskij, *Integrable Lotka–Volterra Systems*, Regul. Chaotic Dyn. 13 (2008) 543–556.

Note that H is not a constant of motion for the deformed system, generally.

A particular deformations is obtained using

$$I_1 = C_1 + g\alpha(x_4), I_2 = C_2 + g\beta(x_4), I_3 = C_3,$$

namely

$$\begin{aligned}\dot{x}_1 &= x_1x_2 + gx_3x_4\alpha'(x_4), \quad \dot{x}_2 = x_2(x_3 - x_1), \\ \dot{x}_3 &= x_3(x_4 - x_2) + gx_3x_4\beta'(x_4), \quad \dot{x}_4 = -x_3x_4.\end{aligned}$$

Therefore, if $\alpha(x_4) + \beta(x_4) = \text{constant}$, then the constant of motion $H = x_1 + x_2 + x_3 + x_4$ of the initial system is also a constant of motion of the deformed system.

Thank you! :)